

Automatic Omniscop e Calibration Using Redundant

**SwFRT, Fermilab,
2009**

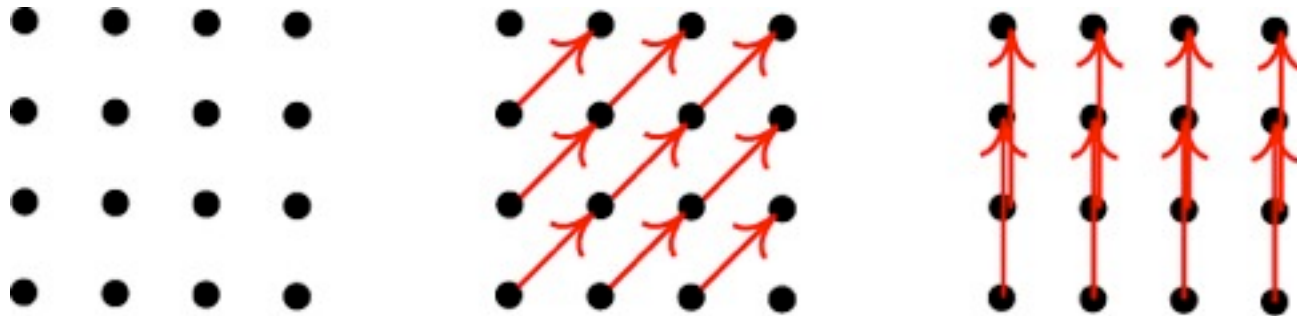
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**(with Max
Tegmark, Andy
Lutomirski, Scott
Morrison, Matias
Zaldarriaga)**

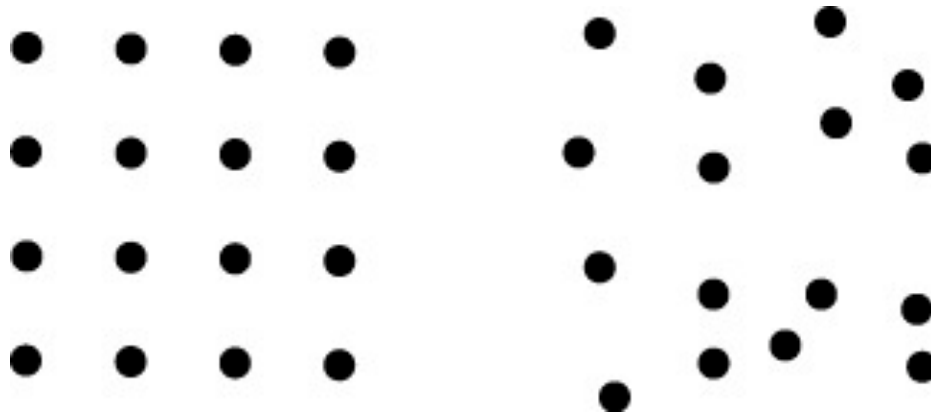


Saturday, October 10, 2009

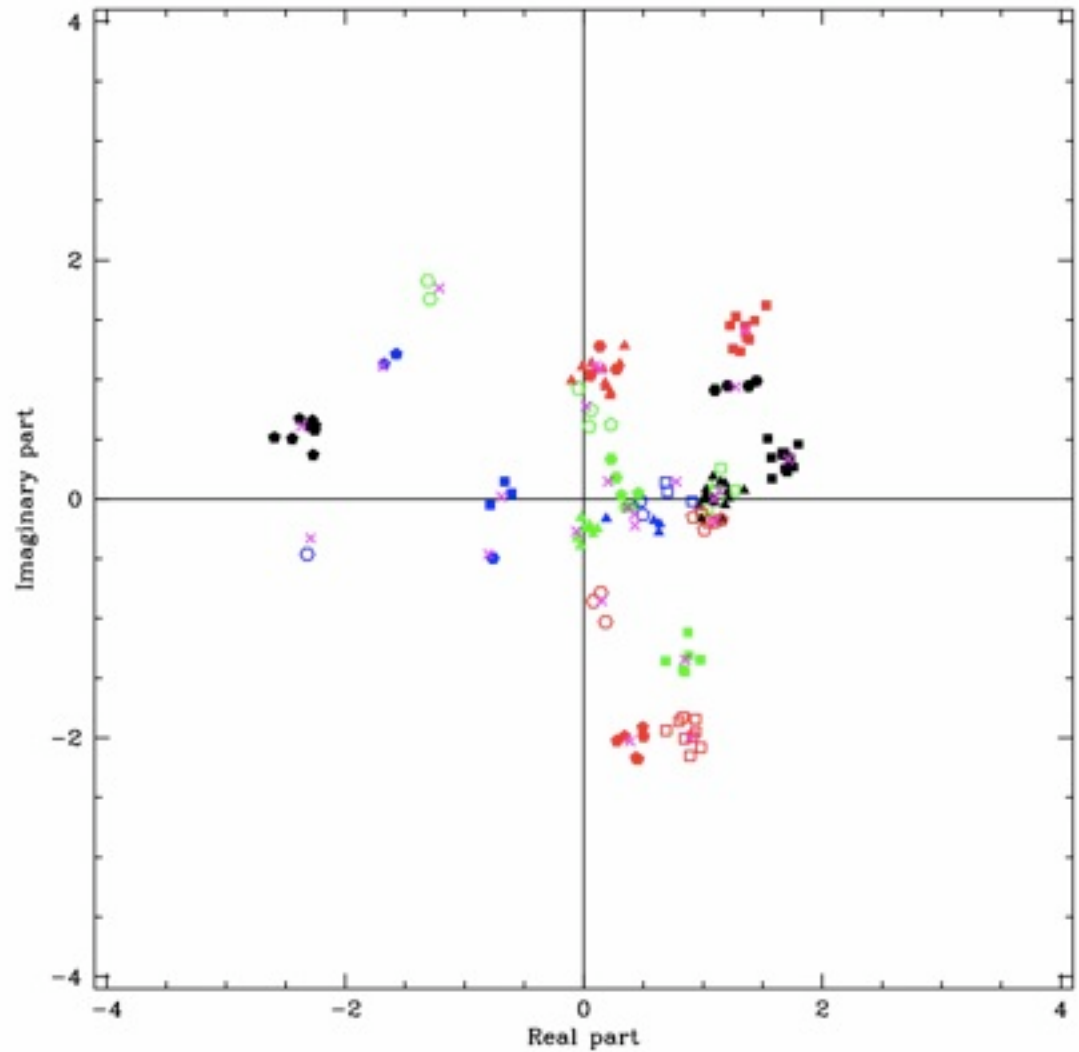
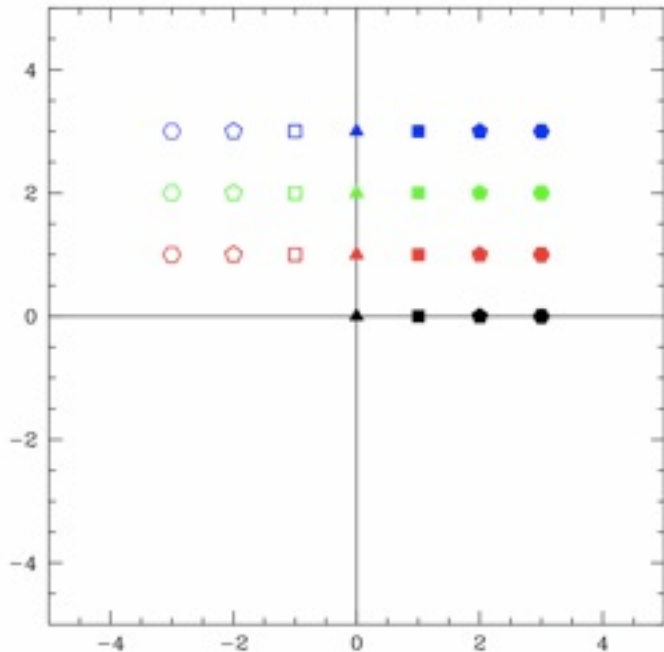
Unlike Traditional

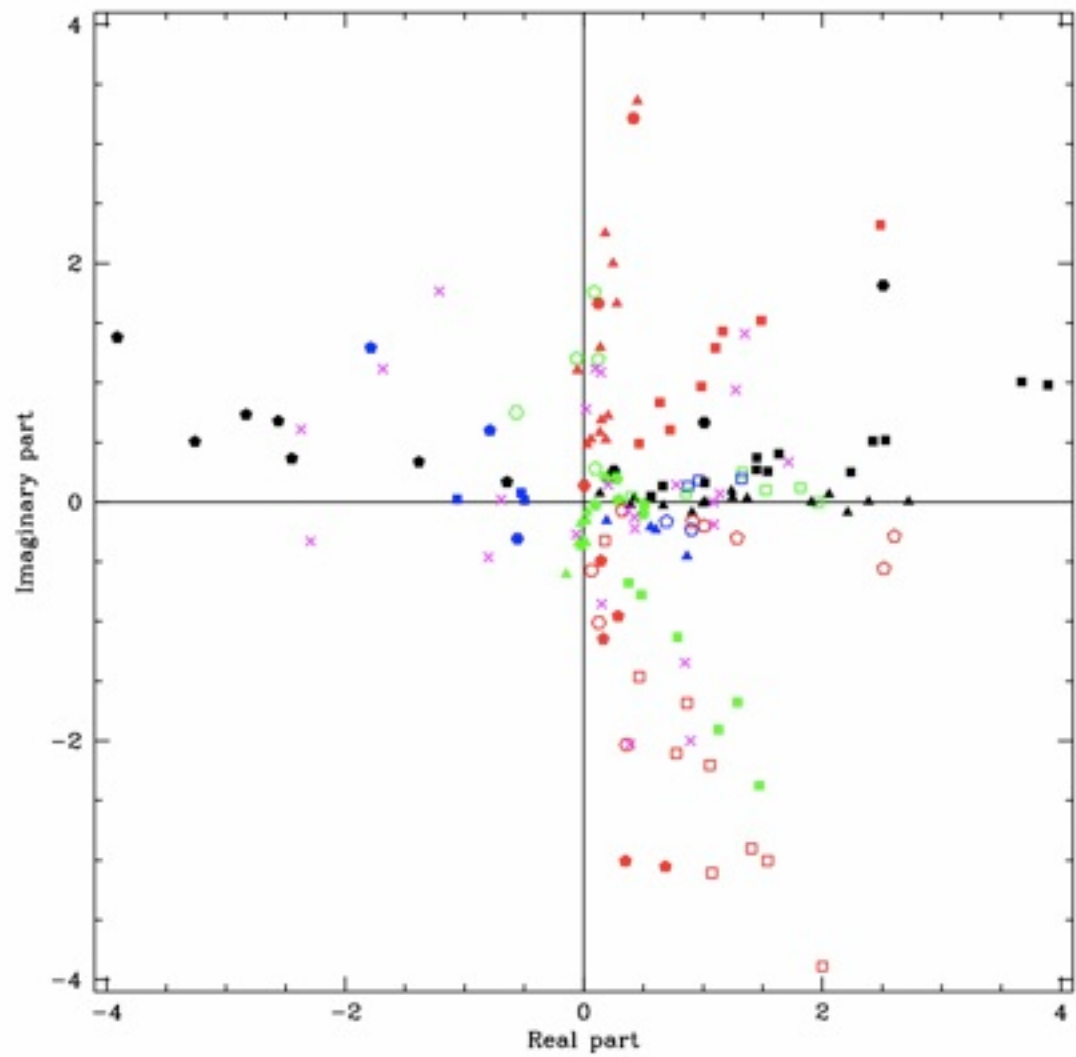
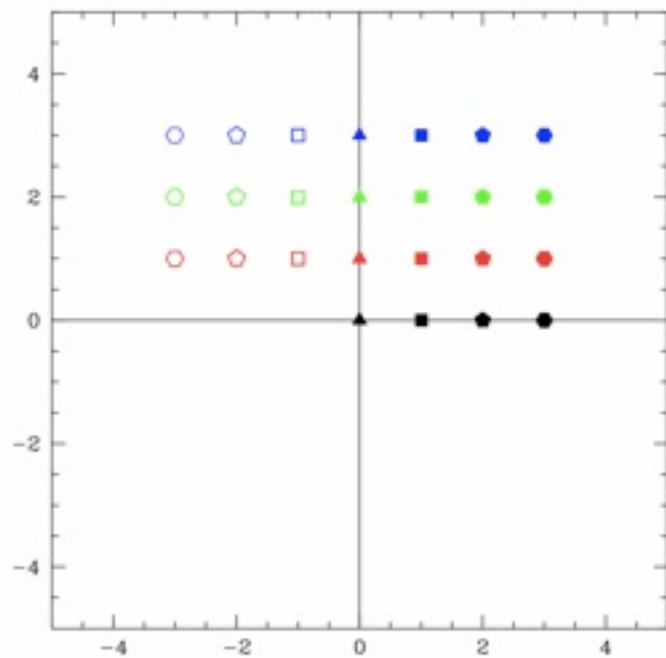


In Reality, Identical Baselines



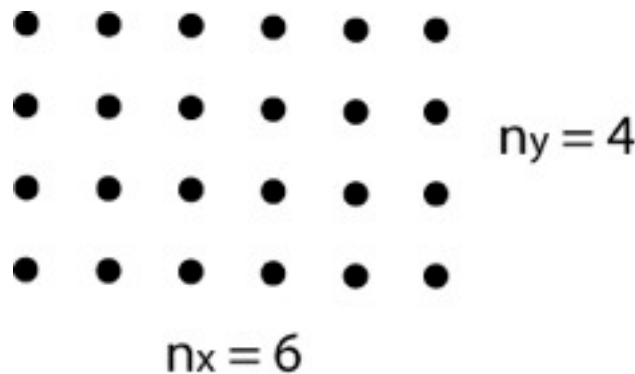
Main Idea:
Demand that
redundant
baselines give
the same
results and fit





Mathematically, an over-determined system

Inputs	Parameters being fit
<ul style="list-style-type: none">• $n_x n_y (n_x n_y + 1) / 2$ complex correlations	<ul style="list-style-type: none">• $n_x n_y$ complex gains• $2 n_x n_y - n_x - n_y + 1$ unique baselines



Measured signal

True sky signal

Instrumental noise

$$s_i(t) = g_i x_i(t) + n_i(t)$$

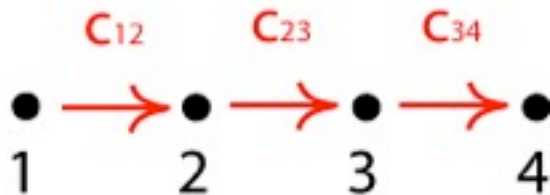
Complex gain:

$$g_i \equiv e^{\eta_i + i\varphi_i}$$

$$\begin{aligned} c_{ij}(t) &\equiv s_i^*(t) s_j(t) \\ &= g_i^* g_j x_i^*(t) x_j(t) + \underbrace{n_i^*(t) n_j(t) + x_i^*(t) n_j(t) + x_j^*(t) n_i(t)}_{\text{Assume zero mean}} \end{aligned}$$

$$\langle c_{ij} \rangle = g_i^* g_j y_{i-j}(t) \quad \text{where} \quad y_{i-j}(t) \equiv \langle x_i^*(t) x_j(t) \rangle$$

s_i = signal measured by antenna i ; g_i = gain of antenna i
 $g_i \equiv e^{\eta_i + i\varphi_i}$
 corr; c_{ii} = measured corr



$$\langle c_{12} \rangle = g_1^* g_2 y_1$$

$$\langle c_{23} \rangle = g_2^* g_3 y_1$$

$$\langle c_{34} \rangle = g_3^* g_4 y_1$$

$$\vdots$$

$$\langle c_{14} \rangle = g_1^* g_4 y_3$$

$$\ln \langle c_{12} \rangle = (\eta_1 + \eta_2 + \ln |y_1|) + i(-\varphi_1 + \varphi_2 + \arg y_1)$$

$$\ln \langle c_{23} \rangle = (\eta_2 + \eta_3 + \ln |y_1|) + i(-\varphi_2 + \varphi_3 + \arg y_1)$$

$$\ln \langle c_{34} \rangle = (\eta_1 + \eta_2 + \ln |y_1|) + i(-\varphi_1 + \varphi_2 + \arg y_1)$$

$$\vdots$$

$$\ln \langle c_{14} \rangle = (\eta_1 + \eta_4 + \ln |y_3|) + i(-\varphi_1 + \varphi_4 + \arg y_3)$$

s_i = signal measured by antenna i ; g_i = gain of antenna i

; $g_i \equiv e^{\eta_i + i\varphi_i}$
corr; c_{ii} = measured corr

Real Part: Gains


$$\begin{pmatrix} \text{Re ln}\langle c_{11} \rangle \\ \text{Re ln}\langle c_{22} \rangle \\ \text{Re ln}\langle c_{33} \rangle \\ \text{Re ln}\langle c_{44} \rangle \\ \text{Re ln}\langle c_{12} \rangle \\ \text{Re ln}\langle c_{23} \rangle \\ \text{Re ln}\langle c_{34} \rangle \\ \text{Re ln}\langle c_{13} \rangle \\ \text{Re ln}\langle c_{24} \rangle \\ \text{Re ln}\langle c_{14} \rangle \end{pmatrix} = \begin{pmatrix} 2 & 0 & 0 & 0 & 1 & 0 & 0 & 0 \\ 0 & 2 & 0 & 0 & 1 & 0 & 0 & 0 \\ 0 & 0 & 2 & 0 & 1 & 0 & 0 & 0 \\ 0 & 0 & 0 & 2 & 1 & 0 & 0 & 0 \\ 1 & 1 & 0 & 0 & 0 & 1 & 0 & 0 \\ 0 & 1 & 1 & 0 & 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 1 & 0 & 1 & 0 & 0 \\ 1 & 0 & 1 & 0 & 0 & 0 & 1 & 0 \\ 0 & 1 & 0 & 1 & 0 & 0 & 1 & 0 \\ 1 & 0 & 0 & 1 & 0 & 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} \eta_1 \\ \eta_2 \\ \eta_3 \\ \eta_4 \\ \ln |y_0| \\ \ln |y_1| \\ \ln |y_2| \\ \ln |y_3| \end{pmatrix}$$

Have this
(data)
Know this
Want this


s_i = signal measured by antenna i ; g_i = gain of antenna i
 $g_i \equiv e^{\eta_i + i\varphi_i}$
 c_{ii} = measured corr

Real Part: Gains


$$\begin{pmatrix} \text{Re ln}\langle c_{11} \rangle \\ \text{Re ln}\langle c_{22} \rangle \\ \text{Re ln}\langle c_{33} \rangle \\ \text{Re ln}\langle c_{44} \rangle \\ \text{Re ln}\langle c_{12} \rangle \\ \text{Re ln}\langle c_{23} \rangle \\ \text{Re ln}\langle c_{34} \rangle \\ \text{Re ln}\langle c_{13} \rangle \\ \text{Re ln}\langle c_{24} \rangle \\ \text{Re ln}\langle c_{14} \rangle \end{pmatrix} = \begin{pmatrix} 2 & 0 & 0 & 0 & 1 & 0 & 0 & 0 \\ 0 & 2 & 0 & 0 & 1 & 0 & 0 & 0 \\ 0 & 0 & 2 & 0 & 1 & 0 & 0 & 0 \\ 0 & 0 & 0 & 2 & 1 & 0 & 0 & 0 \\ 1 & 1 & 0 & 0 & 0 & 1 & 0 & 0 \\ 0 & 1 & 1 & 0 & 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 1 & 0 & 1 & 0 & 0 \\ 1 & 0 & 1 & 0 & 0 & 0 & 1 & 0 \\ 0 & 1 & 0 & 1 & 0 & 0 & 1 & 0 \\ 1 & 0 & 0 & 1 & 0 & 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} \eta_1 \\ \eta_2 \\ \eta_3 \\ \eta_4 \\ \ln |y_0| \\ \ln |y_1| \\ \ln |y_2| \\ \ln |y_3| \end{pmatrix}$$



d



A



b

$$\hat{b} = [A^t A]^{-1} A^t \vec{d}$$

s_i = signal measured by antenna i ; g_i = gain of antenna i

$g_i \equiv e^{\eta_i + i\varphi_i}$; c_{ii} = measured corr

A Problem With Degeneracies

$$\hat{b} = [A^t A]^{-1} A^t \vec{d}$$

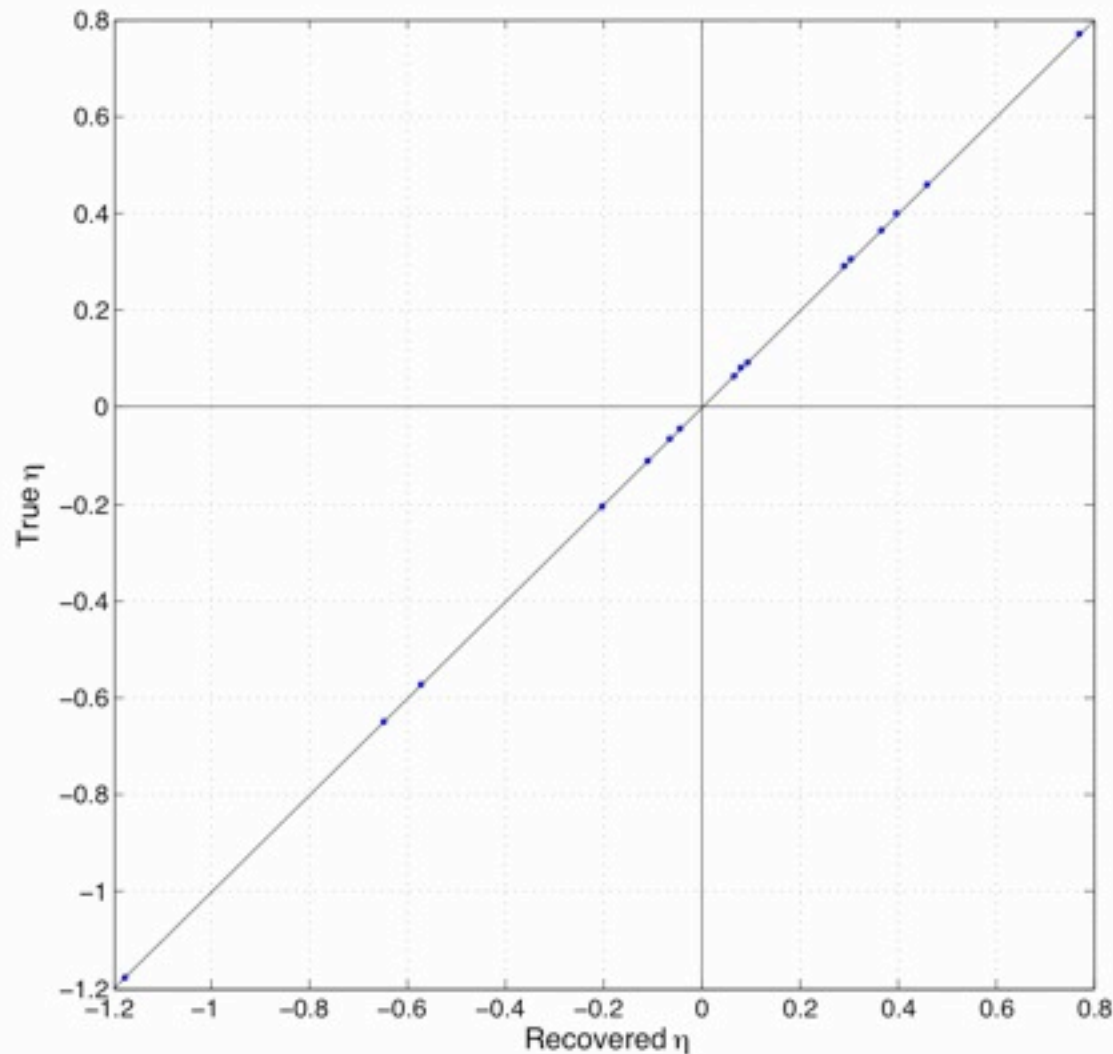
Real part (gains):

- No knowledge of absolute gain.

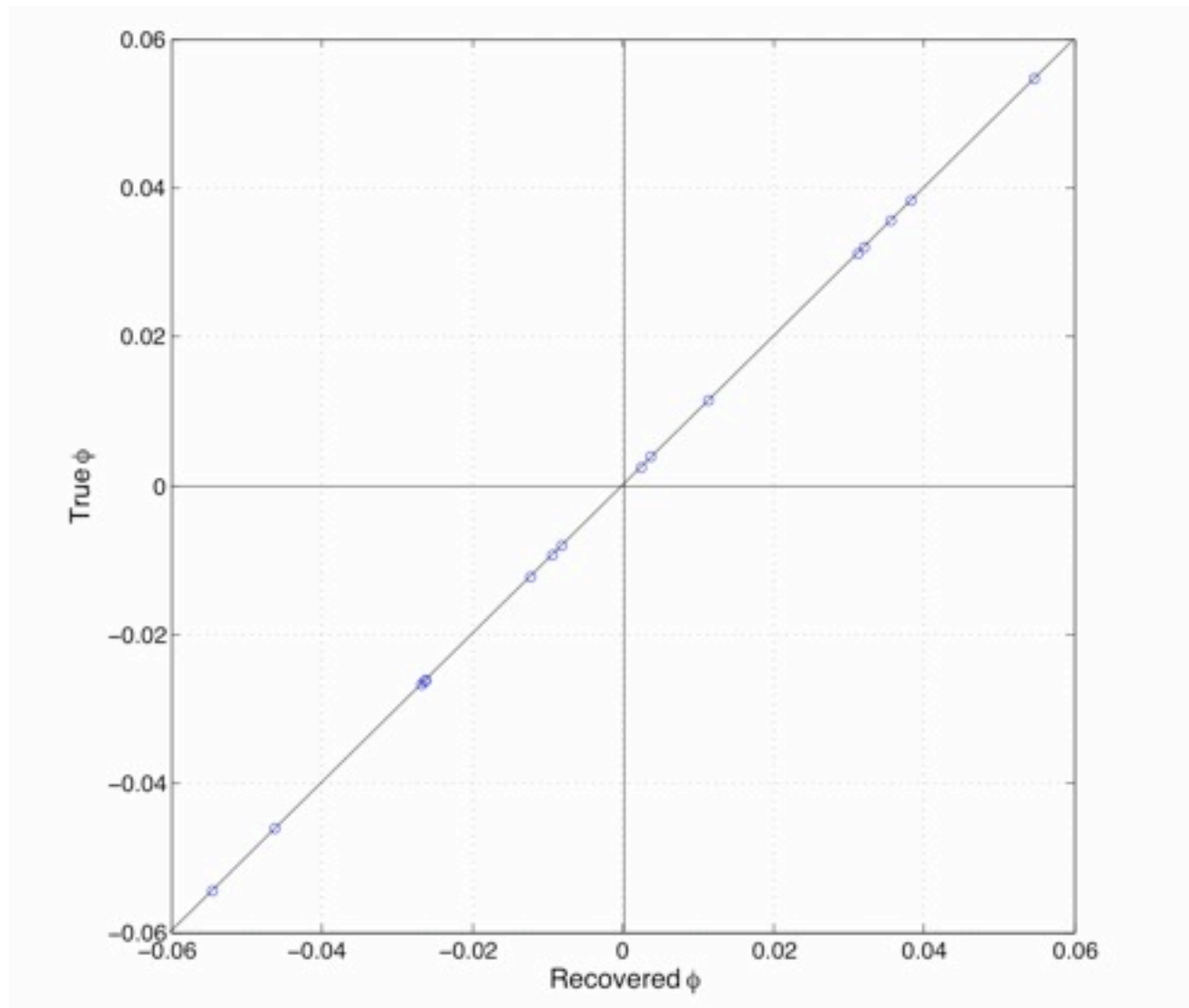
Imaginary part (phases):

- No knowledge of absolute phase.
- No knowledge of x-tilt.
- No knowledge of y-tilt.

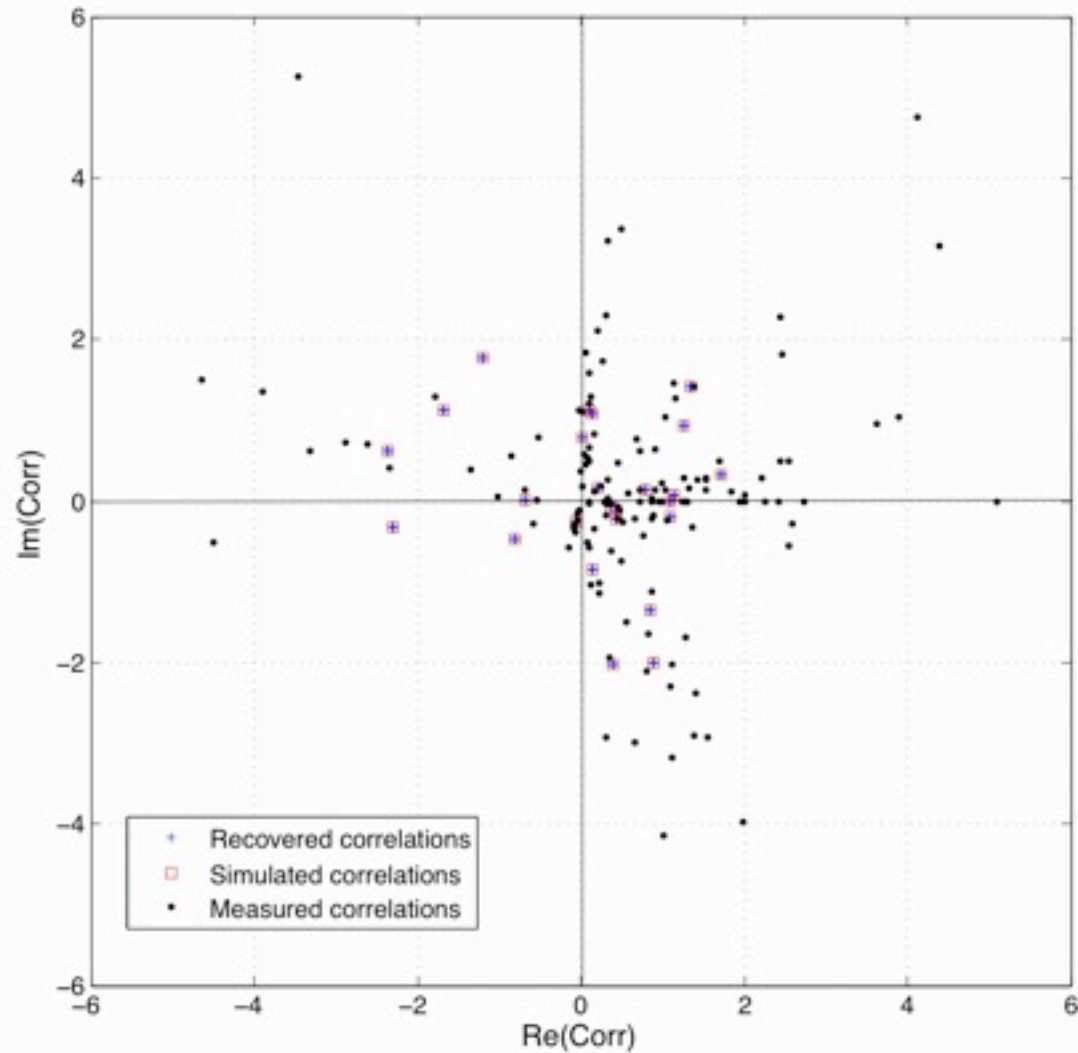
s_i = signal measured by antenna i ; g_i = gain of antenna i
 $g_i \equiv e^{\eta_i + i\varphi_i}$ corr; c_{ii} = measured corr



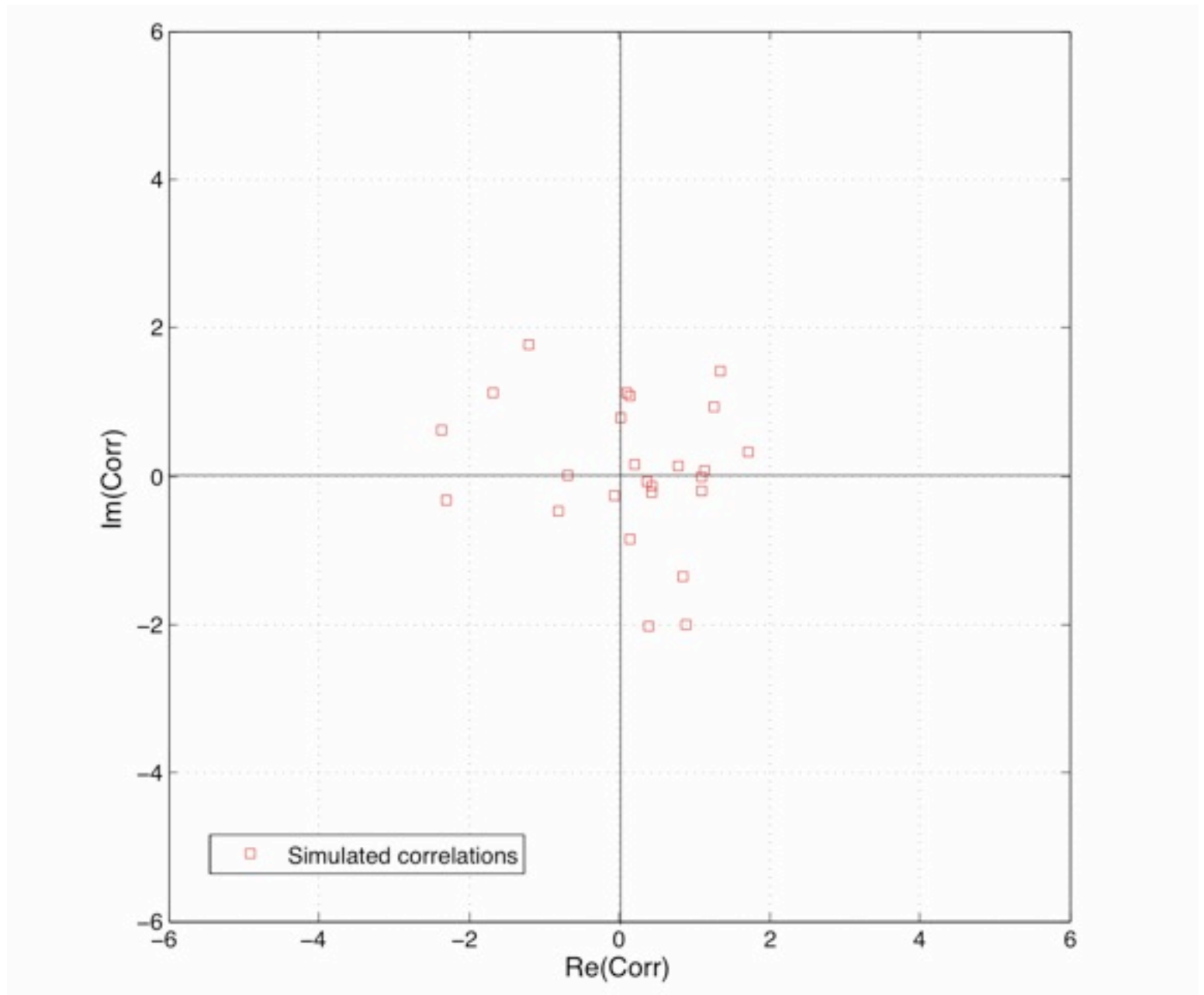
s_i = signal measured by antenna i ; g_i = gain of antenna i
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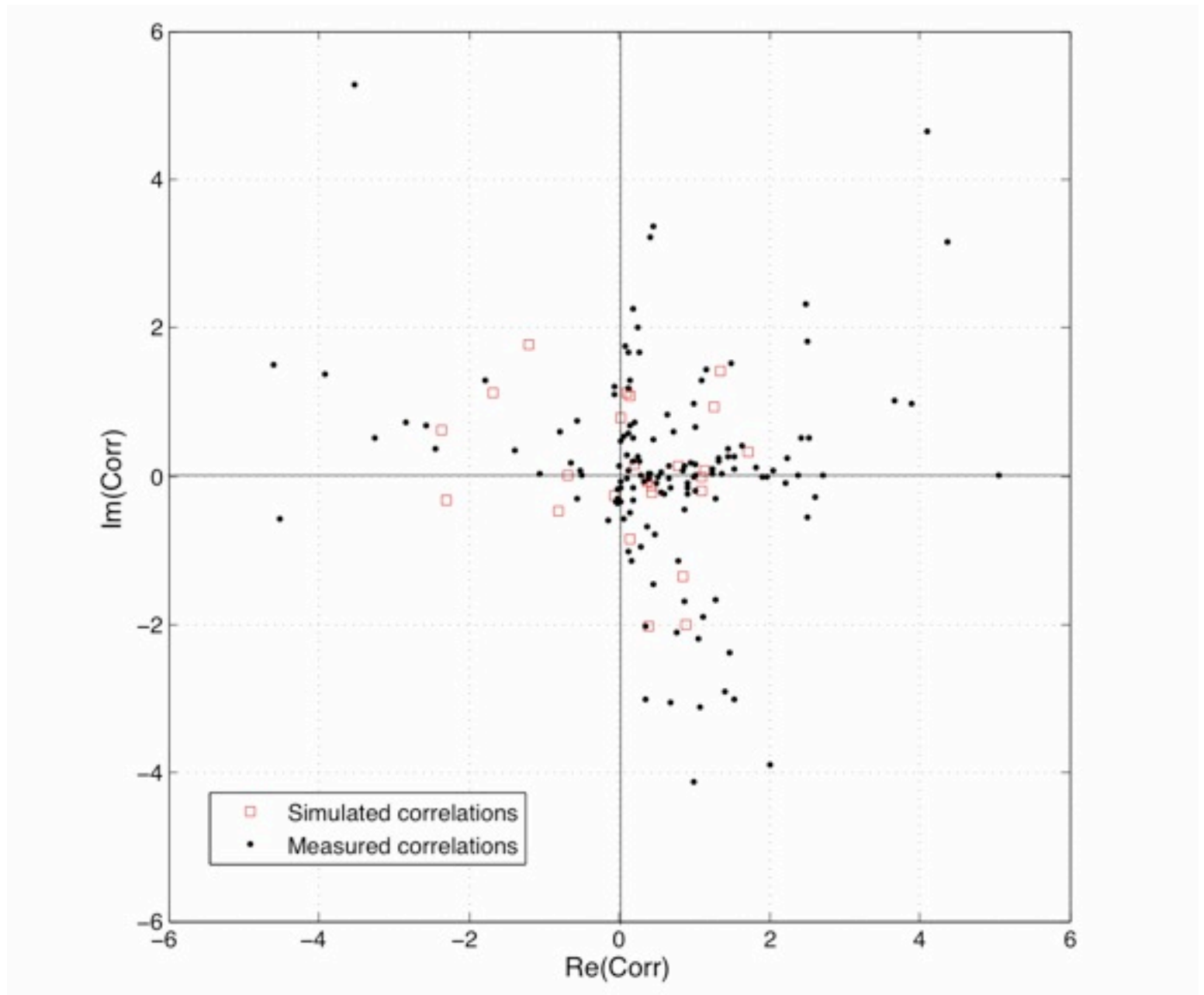


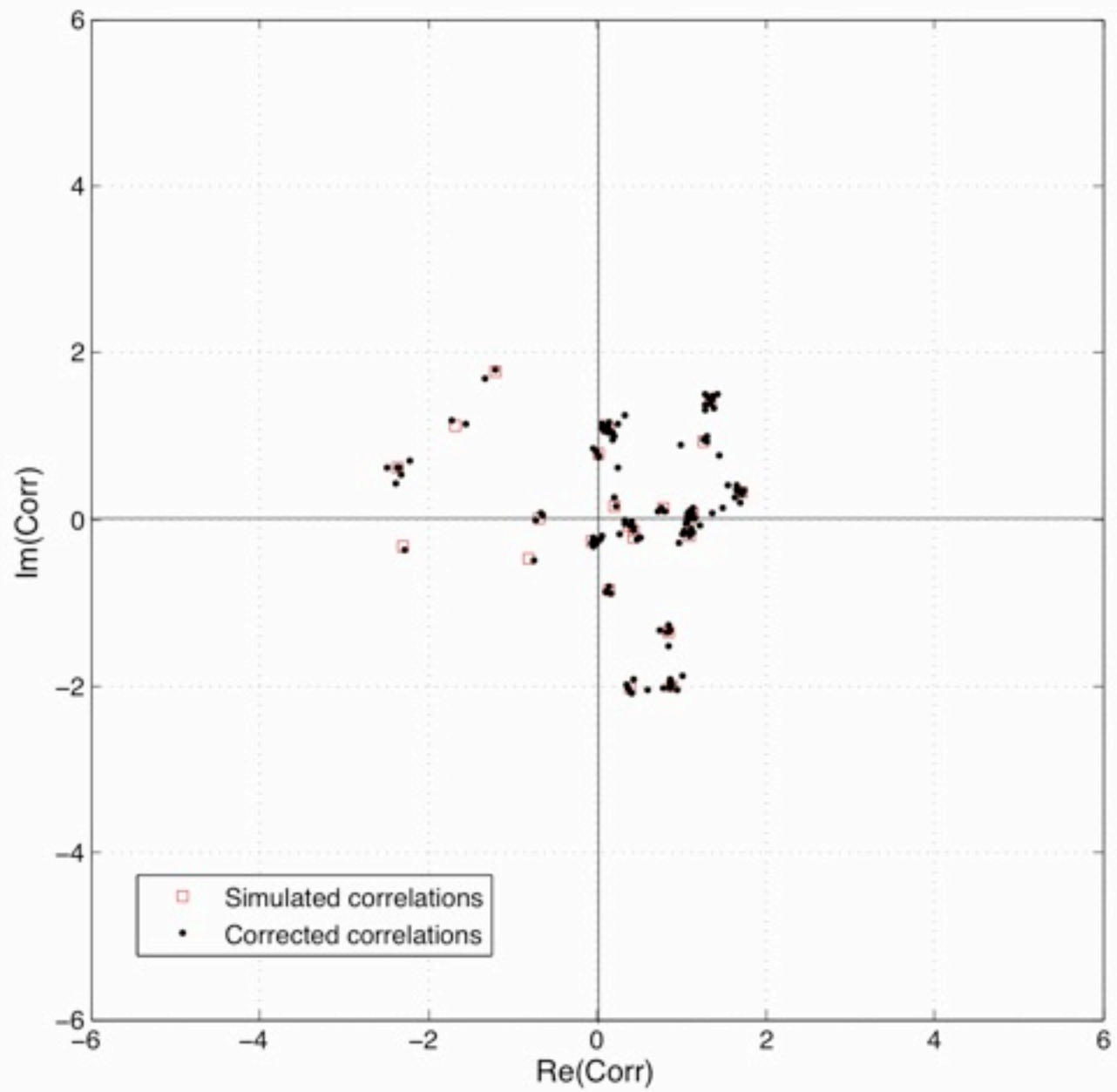
s_i = signal measured by antenna i ; g_i = gain of antenna i
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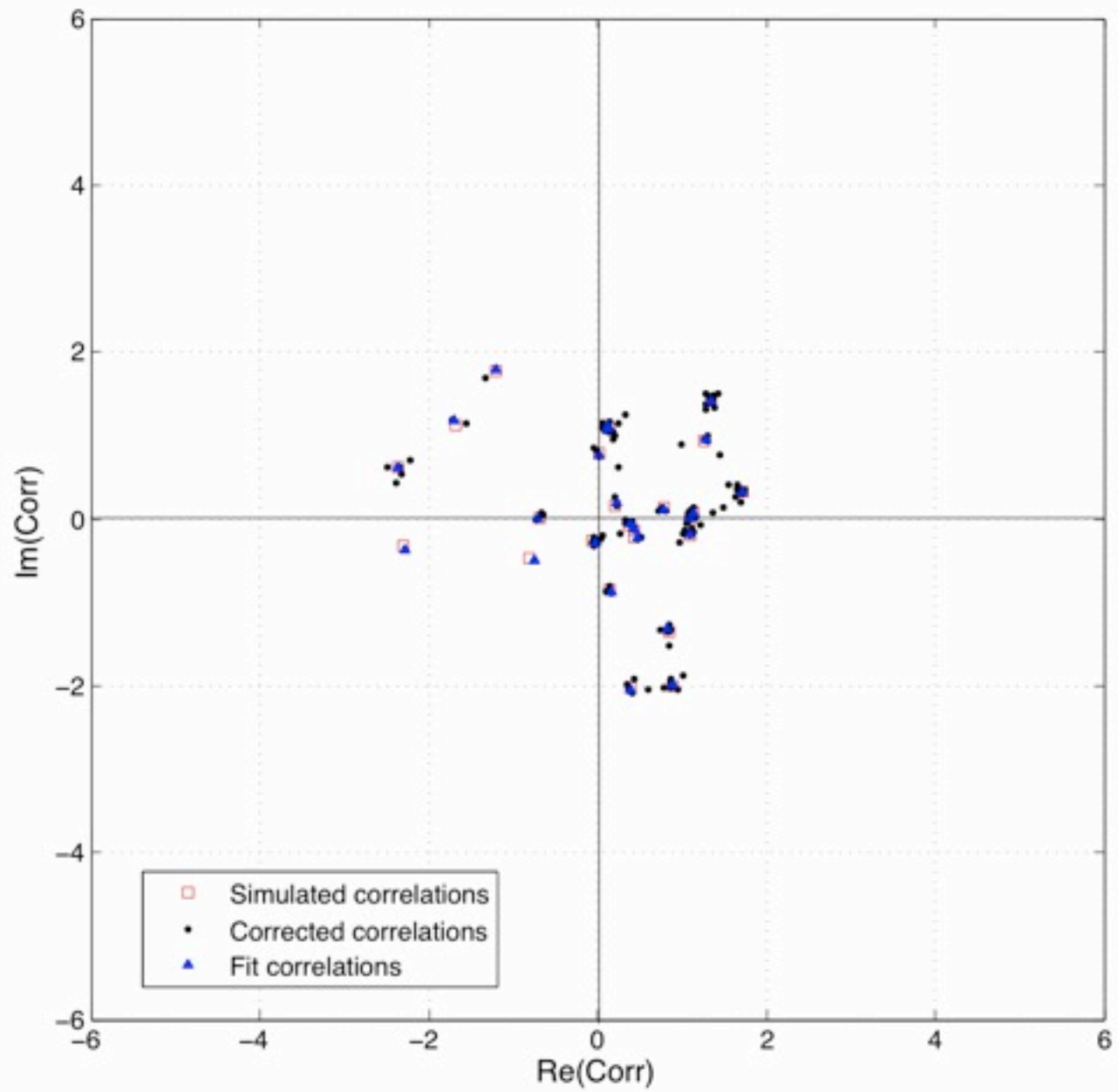


s_i = signal measured by antenna i ; g_i = gain of antenna i
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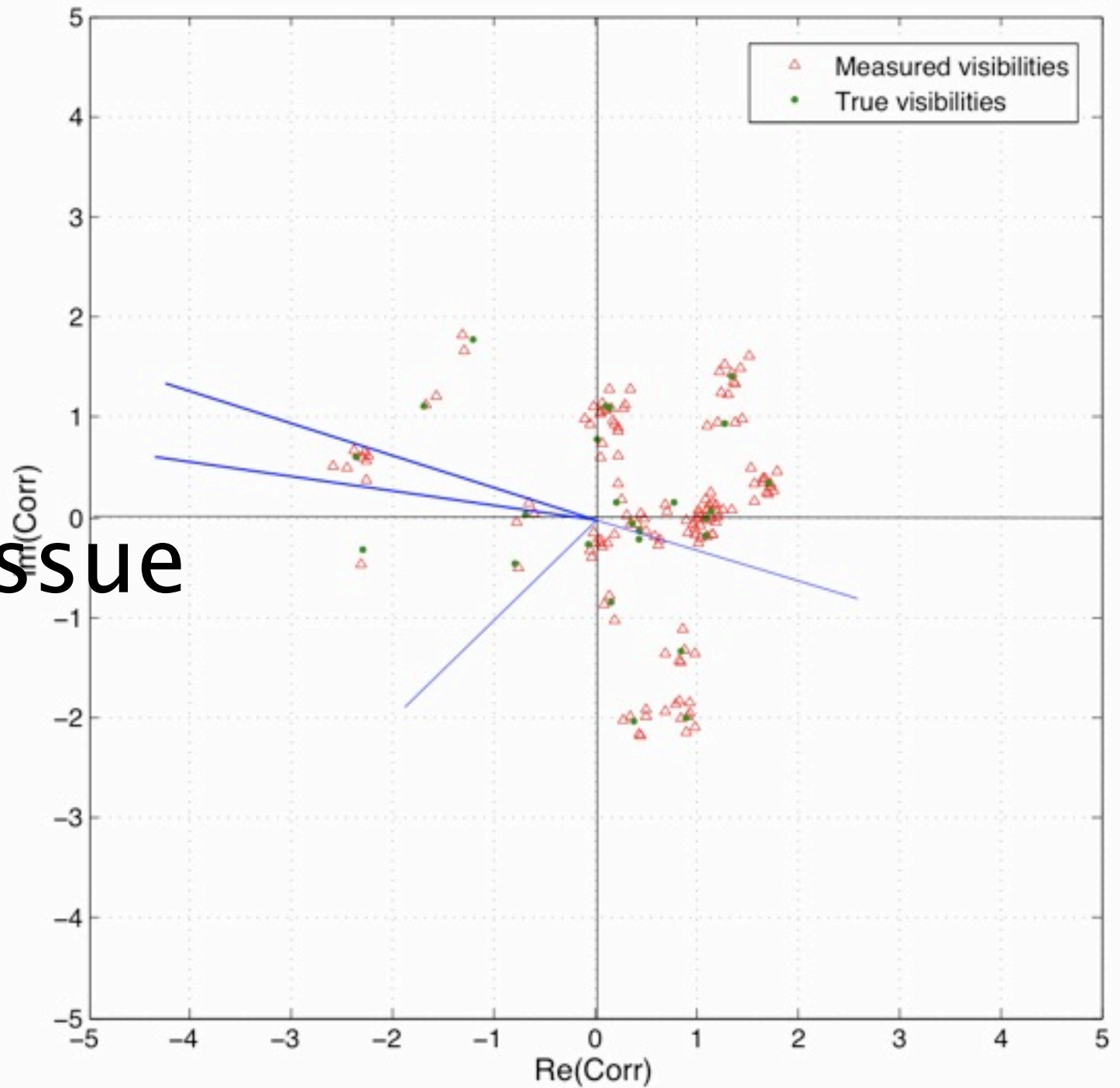








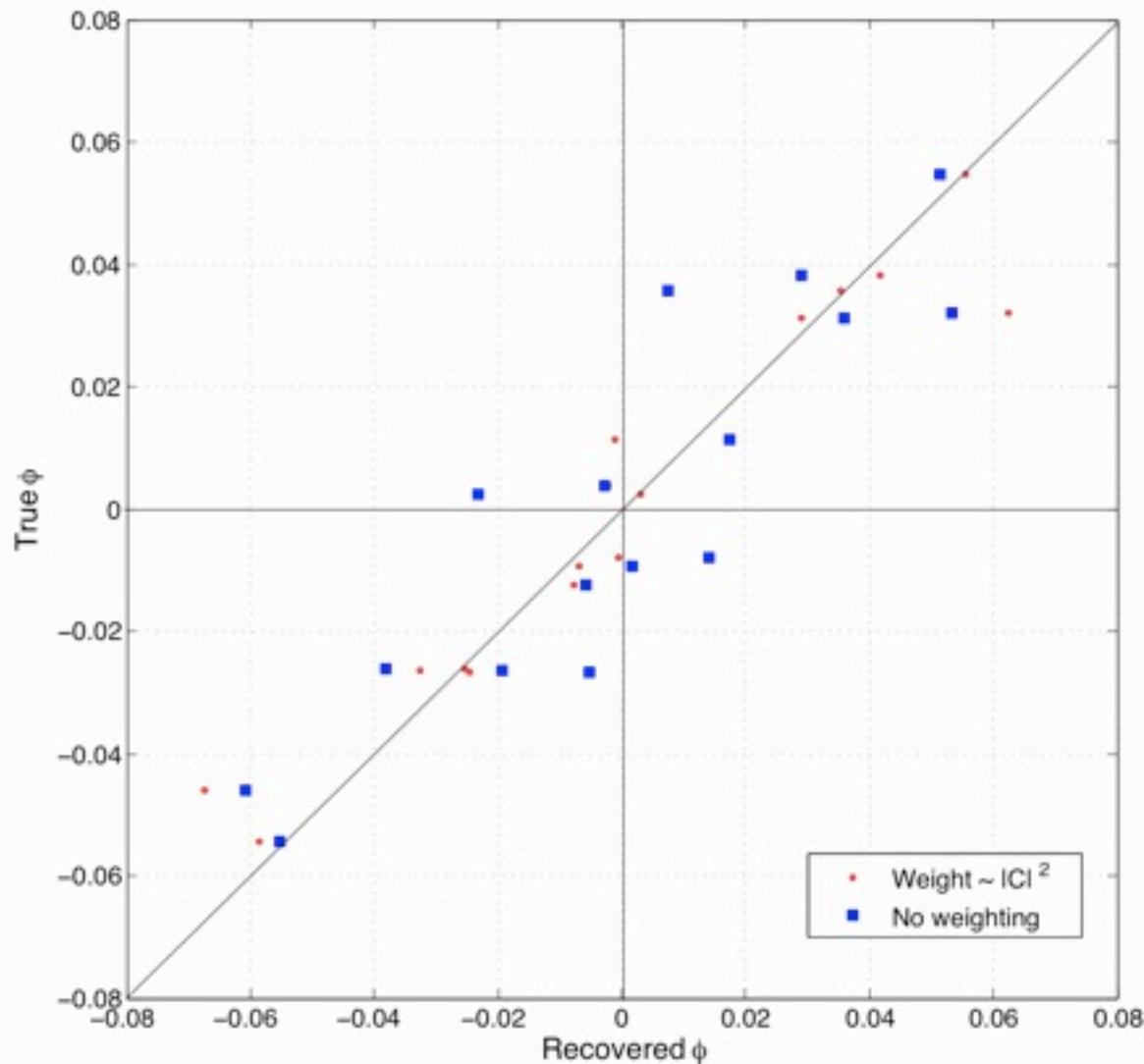
Eighty Issue



Weighted Fit

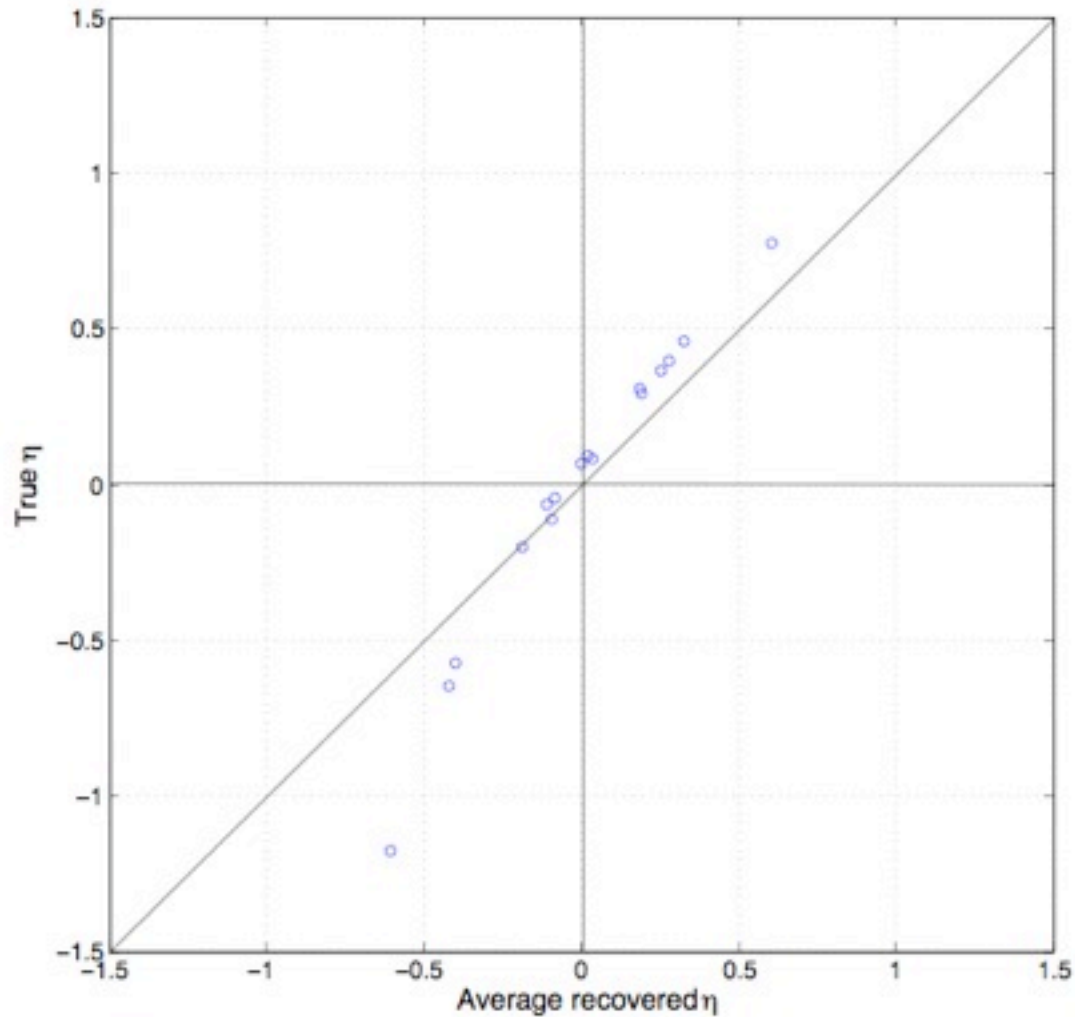
$$\hat{b} = [A^t W A]^{-1} A^t W \vec{d}$$

$$W_{ij} = \delta_{ij} |c_i|^2$$



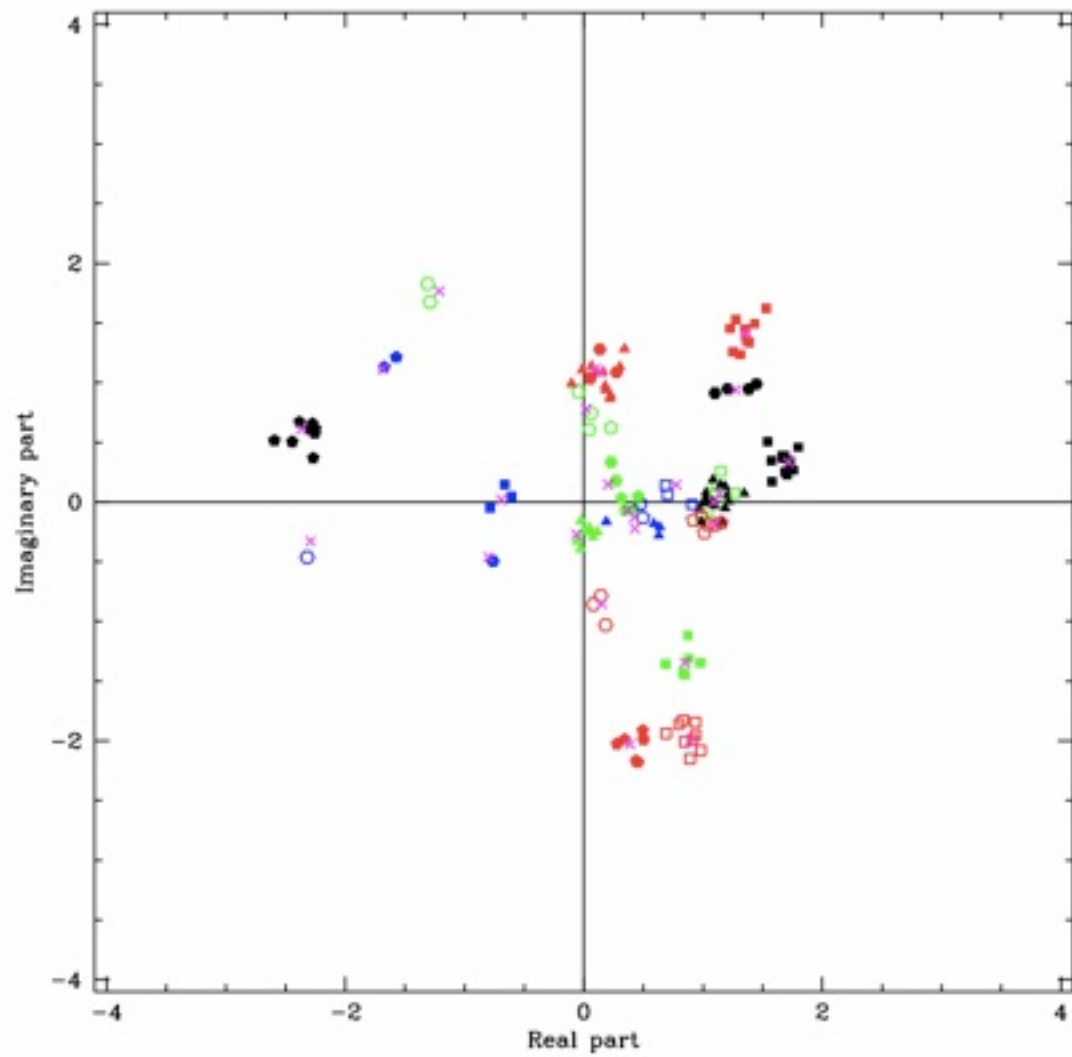
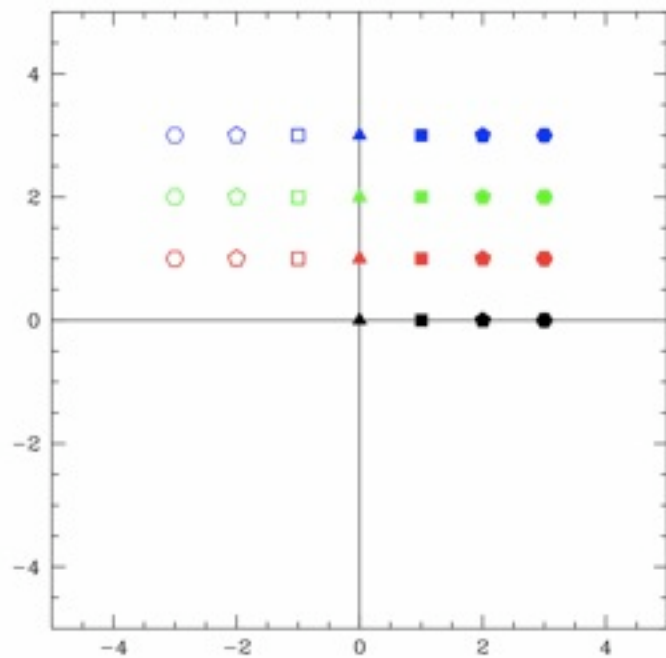
s_i = signal measured by antenna i ; g_i = gain of antenna i
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Some Bias



Differences Between This and Traditional Calibration

Traditional Scheme	Redundant Baselines
<ul style="list-style-type: none">• Use closure phases• Amplitude (gain) calibration requires point source• Modeling may be required	<ul style="list-style-type: none">• Any sky signal can be used for calibration• Redundancy provided by identical baselines



Summary

- Redundant baselines provided by omniscopes can be exploited for calibration
- Proposed automatic calibration scheme works on simulated data
- Using redundant baselines allows one to sidestep certain problems with traditional techniques (like the necessity of having isolated point sources).